Wavefront sensing from speckle images with polychromatic phase diversity

Xavier Rondeau & Eric Thiébaut

Airi Team

Centre de Recherche Astronomique de Lyon

ADA-5th conference

Thursday, May 8th, 2008



High angular resolution in Astronomy

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight

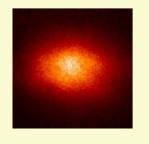
Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

 $\rightarrow \propto \lambda/D$ for an isolated telescope **Angular** resolution $\Rightarrow \propto \lambda/r_0$ with atmospheric turbulence

 $D/r_0(\lambda)\gg 1$





BOA-Onera

▶ $D \simeq 30,40 \,\mathrm{m}$ (TMT,E-ELT), $\lambda = 500 \,\mathrm{nm}, \, r_0(\lambda) = 15 \,\mathrm{cm}$

 \Rightarrow Resolution loss factor D/r₀ $\simeq 200$



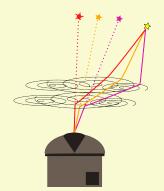
Context HRA in Astronomy P-Laser Guide Star Tip-tilt estimation Phase retrieval pb

Polychromatic Laser Guide Star

R. Fov et al. 1995

tip-tilt indetermination

AO limitations due to



Effectiveness Polychromatic

Conclusion

Inverse problem Bayesian formulation

Global optimization Global strategy Regularisation weight Local optimization

Difficulties

diversity Polychromatic model Results

★ Use of air refractive index chromaticity

Searched tip-tilt
$$\theta(\lambda) = \underbrace{\frac{n(\lambda) - 1}{n(\lambda_2) - n(\lambda_1)}}_{C(\lambda_1, \lambda_2)} \times \underbrace{\left[\theta(\lambda_2) - \theta(\lambda_1)\right]}_{Measured tip-tilts}$$



Tip-tilt estimation

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight Local optimization Effectiveness

Polychromatic diversity Polychromatic model Results

Conclusion

Amplification of the tip-tilt error :

$$\sigma_{\theta_{\lambda}} = \zeta(\lambda_1, \lambda_2) \times \sigma_{\Delta \theta_{\lambda_1, \lambda_2}}$$
 where $\zeta(\lambda_1, \lambda_2) > 25$

Degradation of Cramér-Rao lower bound :

$$\tilde{\theta}_{\mathrm{ML}} = \max_{\theta} \, \mathbf{g}(\mathbf{x}|\theta) \implies \sigma_{\theta}^{+2} = -\left(E\left[\tfrac{\hat{c}^2 \ln(\mathbf{g}(\mathbf{x}|\theta))}{\hat{c}\theta^2}\right]\right)_{\left|\theta = \theta^+\right|}^{-1}$$

Airy pdf:
$$\sigma_{\theta}^{+} = \frac{\kappa}{\sqrt{N_{\mathrm{ph}}}} \times \frac{\lambda}{D}$$

Average pdf of speckle images : $\sigma_{\theta}^{+} \simeq \frac{0.3 \, \kappa}{\sqrt{N_{\mathrm{ph}}}} \times \frac{\lambda}{r_0}$

 $\sim 0.3 \times D/\textit{r}_0 \text{ increase}$ due to turbulence

Joint wavefront estimation from the images to attain the ultimate centering accuracy



The inverse problem

Context HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight

Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model

Results Conclusion model noise

parameters

 $\mathbf{h} = \mathbf{m}(\boldsymbol{\varphi}) + \mathbf{n}$

Forward model:

data

turbulent $\varphi(\mathbf{u})$ observed psf $h_{\lambda,t}(\mathbf{x})$ Telescope pupil

Telescope focal plane

where $\mathbf{m}(\boldsymbol{\varphi}) = \alpha |FT[\mathbf{P} \times \exp(i\boldsymbol{\varphi})]|^2$

The inverse approach : Reconstruction of $\varphi(\mathbf{u})$ given data $h_{\lambda,t}(\mathbf{x})$, forward model $\mathbf{m}(\varphi)$ and statistics of noise \mathbf{n}



Bayesian formulation

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval ph Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy

Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

 $\varphi^{+} = \max_{\boldsymbol{\varphi}} \left\{ \Pr\left[\boldsymbol{\varphi}|\mathbf{h}\right] \right\} = \max_{\boldsymbol{\varphi}} \left\{ \Pr\left[\mathbf{h}|\boldsymbol{\varphi}\right] \times \Pr\left[\boldsymbol{\varphi}\right] \right\}$ MAP criterion

Gaussian prior statistics
$$f_{\mathrm{prior}}(oldsymbol{arphi}) = oldsymbol{arphi}^ op. \mathbf{C}_{oldsymbol{arphi}}^{-1}.oldsymbol{arphi}$$
 where $\mathbf{c}_{oldsymbol{arphi} = \langle oldsymbol{arphi}, oldsymbol{arphi}^ op \rangle}$ (Kolmogorov covariance)

Uncorrelated low count data
$$f_{\exp}^{ ext{Poisson}}(oldsymbol{arphi}) = \sum\limits_{j} [\mathbf{m}(oldsymbol{arphi}) - \mathbf{h} \log(\mathbf{m}(oldsymbol{arphi}))]_{j}$$

Higher count or $f_{\text{exp.}}^{\text{Gauss}}(\varphi) = [\mathbf{h} - \mathbf{m}(\varphi)]^{\top} . \mathbf{C}_{\mathbf{n}}^{-1} . [\mathbf{h} - \mathbf{m}(\varphi)]$ additive gaussian noise





Ill-posed inverse problem

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

- Intrinsic degeneracies of $\mathbf{m}(oldsymbol{arphi})$
 - ightarrow sign of all even phase modes : $m\left[oldsymbol{arphi}(\mathbf{r})
 ight] = m\left[-oldsymbol{arphi}(-\mathbf{r})
 ight]$
 - ightarrow modulo 2π periodicity : $m[\varphi] = m[\varphi + 2\mathbf{k}(\mathbf{r})\pi] \quad \forall \varphi_j$
- lacktriangle Strong local minima of $f_{
 m exp}(oldsymbol{arphi})$ (non-convex criterion)
 - → Translation ambiguity (speckle overlapping between m and h)





- → Loose sign ambiguity of individual even phase modes
- → Alteration of the speckle pattern (new speckles) due to the noise

Need of a global optimization strategy



Large-scale non-linear optimization

Though the criterion is non-convex, faster to make the most of the good local continuity — given size of the pb — than Monte-Carlo

Local quadratic approximation :

$$f(\varphi + \delta \varphi) - f(\varphi) = \mathbf{g}_{\varphi}^{\top} . \delta \varphi + \frac{1}{2} \delta \varphi^{\top} . \mathbf{B}_{\varphi} . \delta \varphi + o(\|\delta \varphi\|^{2})$$

$$\bullet \ \mathbf{g} = \boldsymbol{\nabla} f(\boldsymbol{\varphi}) \quad \text{(gradient)}$$
 \Rightarrow Newton step $\delta \boldsymbol{\varphi}^+ = -\mathbf{B}_{\boldsymbol{\varphi}}^{-1} \cdot \mathbf{g}_{\boldsymbol{\varphi}}$ with $\bullet \ \mathbf{B} = \boldsymbol{\nabla}^2 f(\boldsymbol{\varphi}) \quad \text{(hessian)}$

- $o(\|\delta\varphi\|^2) \simeq 0$
- \mathbf{B}_{φ} too expensive to compute owing to the framerate (~ 50 image/s)
- ▶ \mathbf{B}_{φ} too big for inversion ($N_{\varphi} \times N_{\text{data}} \sim 10^8$ scalars)

Need of fast and very efficient limited-memory local optimization

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval ph Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion







Correlation between values of criterion f_{exp} and wavefront reconstruction error

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization

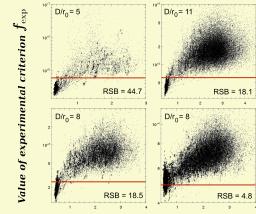
Global strategy

Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

observable



Wavefront reconstruction error (radian rms)







Global optimization strategy

Uncorrelated wavefronts

No proper initialization available for local optimisation but:

- ullet Global convergence guaranteed below some threshold $f_{
 m exp}^{
 m threshold}$
- Decision can be made in very few iterations because of faster convergence rate in the first iterations



> Time-correlated wavefronts

Only local optimization is needed by taking last reconstruction as an initial guess for next wavefront reconstruction (better guess by modeling the wavefront evolution according to Taylor hypothesis)

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight

Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

Tuning of the regularization weight

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy

Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

Weight of convex f_{prior} is enforced in first iterations by tuning

$$\mu \ \ \text{so that} \ \ \delta f_{\exp}[\varphi, \delta \varphi(\mu)] \simeq \epsilon \times \delta f_{\exp}[\varphi, \delta \varphi(\mu=0)]$$

where
$$\delta \boldsymbol{\varphi}(\mu) = -(\mathbf{B}_{\mathrm{exp}} + \mu \, \mathbf{B}_{\mathrm{prior}})^{-1} \, . \, (\mathbf{g}_{\mathrm{exp}} + \mu \, \mathbf{g}_{\mathrm{prior}})$$

- \star constrain the reconstruction to consistent φ by :
 - disentangling $m(\varphi)$ degeneracies
 - smoothing out some local minima
- **★** automatic adaptation to noise and turbulence strength





Local optimization algorithm

HRA in Astronomy P-Laser Guide Star

Context

Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy

Regularisation weight Local optimization

Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

 $N_{\mathrm{dir}} \ll N_{\varphi}$ Subspace of priviledged search directions:

$$\underbrace{ \boldsymbol{\beta}^{+} = -\mathbf{B}'^{-1}.\mathbf{g}' }_{\text{where } \mathbf{B}' = \mathbf{S}^{\top}.\mathbf{B}.\mathbf{S} \text{ and } \mathbf{g}' = \mathbf{S}^{\top}.\mathbf{g} }_{\text{where } \mathbf{B}' = \mathbf{S}^{\top}.\mathbf{B}.\mathbf{S} \text{ and } \mathbf{g}' = \mathbf{S}^{\top}.\mathbf{g}$$



Local optimization algorithm

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval ph Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight

Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion

 $N_{\mathrm{dir}} \ll N_{\varphi}$ Subspace of priviledged search directions:

$$\begin{array}{ccc} \left(\boldsymbol{\beta}^{+} = -\mathbf{B}'^{-1}.\mathbf{g}' \right) & \longleftarrow & \delta f^{\mathrm{quad}}(\boldsymbol{\beta}) = \mathbf{g}'^{\top}.\boldsymbol{\beta} + \frac{1}{2}\,\boldsymbol{\beta}^{\top}.\mathbf{B}'.\boldsymbol{\beta} \\ & \text{where } \mathbf{B}' = \mathbf{S}^{\top}.\mathbf{B}.\mathbf{S} \text{ and } \mathbf{g}' = \mathbf{S}^{\top}.\mathbf{g} \end{array}$$

■ Directions built according to Taylor expansion, using the prior to ensure valid metric with fast fractal implementation (FRIM Thiébaut & Tallon 2008)

•
$$\mathbf{s}_1^{(k)} = \delta \boldsymbol{\varphi}^{(k-1)}$$
 • $\mathbf{s}_2 = -\mathbf{B}_{\mathrm{prior}}^{-1} \cdot \mathbf{g}_{\mathrm{prior}}$

•
$$\mathbf{s}_3 = -\mathbf{B}_{\mathrm{prior}}^{-1} \cdot \mathbf{g}_{\mathrm{exp}}$$
 • $\mathbf{s}_{i+1} = -\mathbf{B}_{\mathrm{prior}}^{-1} \cdot \mathbf{B}_{\mathrm{exp}} \cdot \mathbf{s}_i$ (e.g. $i = 4$,

 $N_{\rm dir} \ll N_{\varphi}$

11



Local optimization algorithm

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval ph Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight Local optimization Effectiveness

Polychromatic

diversity Polychromatic model Results Conclusion

Subspace of priviledged search directions:

 $\delta f^{\mathrm{quad}}(\boldsymbol{\beta}) = \mathbf{g'}^{\top}.\boldsymbol{\beta} + \frac{1}{2}\boldsymbol{\beta}^{\top}.\mathbf{B'}.\boldsymbol{\beta}$ where $\mathbf{B}' = \mathbf{S}^{\mathsf{T}}.\mathbf{B}.\mathbf{S}$ and $\mathbf{g}' = \mathbf{S}^{\mathsf{T}}.\mathbf{g}$

■ $\mathbf{A}_{\mathrm{exp}}$ is an approximation of the too expensive hessian $\mathbf{B}_{\mathrm{exp}}$:

 $\mathbf{A}_{\mathrm{exp}}^{\mathrm{gauss}} = \mathbf{J}^{\top}.\mathbf{C}_{\mathbf{n}}^{-1}.\mathbf{J}$ Linearizing $\mathbf{m}(\varphi)$ yields $\mathbf{A}_{\text{evp}}^{\text{poisson}} = \mathbf{J}^{\top}.\text{Diag}\left[\frac{\mathbf{h}}{\mathbf{m}^2}\right].\mathbf{J}$

> where jacobian $\mathbf{J}_{k,I}=rac{\partial m_k(oldsymbol{arphi})}{\partial \omega_k}$ need not be formed and is implemented for fast vector product

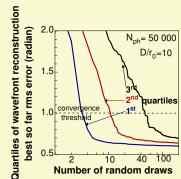
■ Step-length controlled via trust-region to valid quadradic approximation

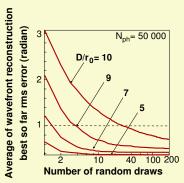
 $\min_{\delta oldsymbol{arphi}} \delta f_{oldsymbol{arphi}}^{\mathrm{quad}}(\delta oldsymbol{arphi}) \;\; ext{so that} \;\; \left\| \delta oldsymbol{arphi}
ight\|_{oldsymbol{\mathsf{M}}}^2 \leqslant \Delta$



Cost with respect to the turbulence strength

Number of random wavefront draws (with prior covariance) needed to attain a given reconstruction error after very few local iterates is investigated through quartiles of a large sampling of wavefront's statistics





High increase of local minima with D/r_0 (bad tip-tilt estimation by image centroïd)

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight

Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results Conclusion





Final convergence

Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization

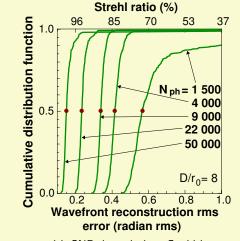
Global strategy Regularisation weight Local optimization

Effectiveness

Polychromatic diversity

Polychromatic model Results

Conclusion



In agreement with SNR degradation. Could be compared to a linearized estimator of the covariance error assuming local convexity

Polychromatic phase diversity

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model

Results Conclusion Use of images at different wavelengths with same optical path

Polychromatic model: $\varphi_{\lambda_i}(\mathbf{r}) = \frac{\lambda_{\text{ref}}}{\lambda_i} \frac{(n_{\lambda_i} - 1)}{(n_{\lambda_{\text{ref}}} - 1)} \times \varphi_{\lambda_{\text{ref}}}(\mathbf{r})$

$$\text{MAP criterion: } f(\varphi_{\lambda_{\mathrm{ref}}}) = \sum_k \mu_k \, f_{\mathrm{exp},\lambda_k}(\varphi_{\lambda_{\mathrm{ref}}}) + f_{\mathrm{prior}}(\varphi_{\lambda_{\mathrm{ref}}})$$

- \star $f_{\exp,\lambda}$ at long λ used as a smooth guide for shorter wavelength:

 (local minima and degeneracies \nearrow when λ \(\)
- μ_1^+ so that $\delta f_{\exp,\lambda_1}[\delta \varphi(\mu_1,\mu_2=0)] \simeq \epsilon_1 \times \delta f_{\exp,\lambda_1}[\delta \varphi(\mu_1=\infty,\mu_2=0)]$ $\stackrel{\downarrow}{\epsilon}_{[0,1]}$ then
- μ_2^+ so that $\delta f_{\exp,\lambda_2}[\delta \varphi(\mu_1^+,\mu_2)] \simeq \epsilon_2 \times \delta f_{\exp,\lambda_2}[\delta \varphi(\mu_1^+,\mu_2=\infty)]$

Polychromatic level-arm enables to deal with much higher D/r_0 (~80 i.e. $N_{\varphi} \ge 7000$)



Results

imulate

S

Context HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

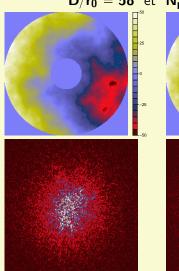
Global optimization Global strategy

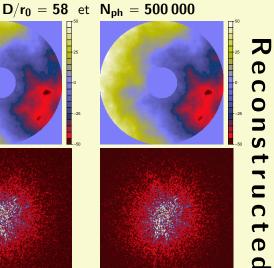
Regularisation weight Local optimization Effectiveness

diversity

Results

Polychromatic Polychromatic model Conclusion







Context HRA in Astronomy

P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization

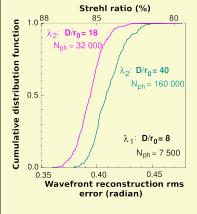
Global strategy Regularisation weight Local optimization Effectiveness

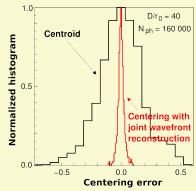
Polychromatic diversity

Polychromatic model Results

Conclusion

Results





Effective ultimate centering accuracy thanks to joint wavefront sensing



Conclusion and perpectives

Rondeau et al. 2007

★ Global optimisation for the wavefront reconstruction pb

- \bullet D/r_0 from 4 (previous works) to 11 with a single wavelength
- \bullet $D/r_0 \simeq 70$ using polychromatic diversity with 2 wavelengths
- Automatic strategy with respect to noise and turbulence fluctuations
- ◆ Fast limited-memory local optimization and global strategy

Application to myopic deconvolution or adaptive optics

 \star Benefit for tip-tilt estimation $\simeq 0.3 \, D/r_0$

Blind deconvolution with an interferometric P-LGS

Context

HRA in Astronomy P-Laser Guide Star Tip-tilt estimation

Phase retrieval pb Inverse problem Bayesian formulation Difficulties

Global optimization Global strategy

Regularisation weight Local optimization Effectiveness

Polychromatic diversity

Polychromatic model Results

Conclusion